

Factors controlling the magnetosphere response to interplanetary shocks

Denny M. Oliveira^{1,2}, James M. Weygand³, Eftyhia Zesta², Chigomezyo M. Ngwira⁴, Michael D. Hartinger⁵, Zhonghua Xu^{6,7,8}, Barbara L. Giles², Dan J. Gershman², Marcos V. D. Silveira¹⁰, Vítor M. Souza¹⁰

¹Goddard Planetary Heliophysics Institute, University of Maryland, Baltimore County, Baltimore, MD, USA

²Heliophysics Science Division, NASA Goddard Space Flight Center, Greenbelt, MD, USA

³Department of Earth, Planetary, and Space Sciences, University of California Los Angeles, Los Angeles, CA, USA

⁴ASTRA LLC, Louisville, CO, USA

⁵Space Science Institute, Boulder, CO, USA

⁶Center for Space Science and Engineering Research, Virginia Tech, Blacksburg, VA, USA

⁷Bradley Department of Electrical and Computer Engineering, Virginia Tech, Blacksburg, VA, USA

⁸National Institute of Aerospace, Hampton, VA, USA

⁹Atmospheric and Space Technology Research Associates, LLC, Boulder, CO, USA

¹⁰National Institute for Space Research – INPE, São José dos Campos, SP, Brazil

Abstract: The impact of interplanetary shocks on the magnetosphere triggers geomagnetic activity that can be measured in the geospace and on the ground, such as variations of energetic particle flux in the radiation belts, geospace magnetic field and wave perturbations, auroral substorms, and ground geomagnetic field variations, which can generate geomagnetically induced currents (GICs) in artificial ground conductors. Moreover, there has been significant evidence in past years that the shock impact angle, the angle between the shock normal vector and the Sun-Earth line, is a major factor controlling the subsequent geomagnetic activity. As shown by many simulation and experimental works, geomagnetic activity occurring under symmetric magnetospheric compressions are usually higher when the compression is asymmetric. This is explained by fast and effective compressions of the most important magnetosphere-ionosphere current systems in case of nearly head-on impacts. In this talk, we briefly present the framework for shock normal determinations and summarize general results. We then focus on two instances of shock geoeffectiveness: (i) a few minutes after shock impact, when the major perturbation sources come from the dayside magnetopause (Chapman-Ferraro) current; and (ii), 30-60 minutes after shock impact, when the major perturbation sources come from the magnetotail due to energetic particle injections occurring during substorm times. The geoeffective response is characterized

by multipoint dB/dt variations in North America and around the world. By comparing all cases, we will show that dB/dt variations are more intense, occur earlier, and cover larger geographic areas as a result of nearly frontal shock impacts. Therefore, we will show that the shock impact angle is a very important feature when predicting geomagnetic activity triggered by shocks that are bound to impact Earth with different orientation.

Acknowledgment: We thank the support of NASA Space Weather Operations 2 Research grant # XXXXX

References:

Oliveira, D. M., & Raeder, J. (2014). Impact angle control of interplanetary shock geoeffectiveness. *Journal of Geophysical Research: Space Physics*, 119(10), 8188-8201.

<https://doi.org/10.1002/2014JA020275>

Oliveira, D. M., & Raeder, J. (2015). Impact angle control of interplanetary shock geoeffectiveness: A statistical study. *Journal of Geophysical Research: Space Physics*, 120(6), 4313-4323.

<https://doi.org/10.1002/2015JA021147>

Oliveira, D. M., & Samsonov, A. A. (2018). Geoeffectiveness of interplanetary shocks controlled by impact angles: A review. *Advances in Space Research*, 61(1), 1-44.

<https://doi.org/10.1016/j.asr.2017.10.006>

Oliveira, D. M., Arel, D., Raeder, J., Zesta, E., Ngwira, C. M., Carter, B. A., Yizengaw, E., Halford, A. J., Tsurutani, B. T., & Gjerloev, J. W. (2018b). Geomagnetically induced currents caused by interplanetary shocks with different impact angles and speeds. *Space Weather*, 16 (6), 636-647.

<https://doi.org/10.1029/2018SW001880>

Session: Space plasma physics and nonlinear processes in space geophysics

Oral or Poster: Oral